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Towards a fundamental theory of nature



Open questions

- What is the origin of space-time?
- Which is the fundamental symmetry of nature?
- How does spontaneous symmetry breaking work?
- Why do we observe three copies of matter states?

• ...

Possible answer

• Nature is supersymmetric!

More than a symmetry

Latin: superus—being above/beyond, part of the Olympus, celestial

Haag-Łopuszánski-Sohnius theorem

- evades no-go theorem by Coleman and Mandula ["All Possible Symmetries of the S Matrix"; Coleman, Mandula 1967]
- only direct products of any internal symmetries and Poincaré group allowed (commutator relations)
- supersymmetry is an exception (anticommutator relations) [Haag,Łopuszánski and Sohnius 1975]

a SUSY algebra obeys a *pseudo* Lie algebra

$$\{Q^N_\alpha, \bar{Q}^M_\beta\} = 2\gamma^\mu_{\alpha\beta} P_\mu \delta^{NM},$$

where N,M count the number of SUSY generators, α,β are (Dirac) spinor indices, μ a Lorentz index, P_μ the 4-momentum and Q,\bar{Q} generators of a supersymmetry

Supersymmetry is there

- more than a solution to the hierarchy problem
- more than simply stabilizing the Higgs mass
- less than "doubling" the number of states



[https://blogphysica.wordpress.com]

How SUSY acts

$$Q|\text{scalar}\rangle = |\text{fermion}\rangle$$

 $\overline{Q}|\text{fermion}\rangle = |\text{scalar}\rangle$
 $QQ|\text{scalar}\rangle = 0$

but

$$Q |\mathsf{fermion'}\rangle = |\mathsf{vector}\rangle$$

at least two different kind of multiplets

The (left) chiral Supermultiplet

 $\Phi = \{\phi, \xi, F\}$ with an auxiliary (scalar) field F

$$\Phi(y,\theta) = \phi(y) + \sqrt{2}\theta\xi(y) + \theta\theta F(y),$$

for superspace coordinates $y^{\mu} = x^{\mu} - i \theta \sigma^{\mu} \bar{\theta}$ and Grassmann numbers θ (two-component spinors)

How SUSY acts

$$Q|\text{scalar}\rangle = |\text{fermion}\rangle$$

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but

$$Q|\mathsf{fermion'}\rangle = |\mathsf{vector}\rangle$$

at least two different kind of multiplets

The vector Supermultiplet

 $V = \{A_{\mu}, \lambda, D\}$ with an auxiliary (scalar) field D

$$\Phi(x,\theta,\bar{\theta}) = \theta \sigma^{\mu} \bar{\theta} A_{\mu}(x) + \theta \theta \bar{\theta} \bar{\lambda}(x) + \theta \lambda(x) \bar{\theta} \bar{\theta} + \frac{1}{2} \theta \theta \bar{\theta} \bar{\theta} D(x),$$

in a special supergauge (Wess–Zumino gauge) and where $\theta_{\alpha}^2 = 0$, but $\theta\theta = \theta^{\alpha}\theta_{\alpha} \neq 0$

SUSY heroes: The Minimal Supersymmetric Standard Model



Chiral Superfields of the MSSM

		spin 0	spin 1/2	$\mathrm{SU}(3)_c, \mathrm{SU}(2)_L, \mathrm{U}(1)_Y$	
	Q	$\left(ilde{u}_L, ilde{d}_L ight)$	(u_L,d_L)	$(3,2,rac{1}{6})$	
	\bar{U}	\tilde{u}_R^*	u_R^c	$(ar{3},1,-rac{2}{3})$	
	\bar{D}	$ ilde{d}_R^*$	d_R^c	$(ar{3},1,rac{1}{3})$	
	L	$(ilde{ u}_L, ilde{e}_L)$	(u_L, e_L)	$(1,2,- frac{1}{2})$	
	Ē	\tilde{e}_R^*	e_R^c	$(\bar{1}, 1, 1)$	
	H_u	$\left(h_{u}^{+},h_{u}^{0}\right)$	$\left(\tilde{h}_{u}^{+}, \tilde{h}_{u}^{0}\right)$	$(1,2,rac{1}{2})$	
	H_d	$\left(h_d^0,h_d^-\right)$	$\left(\tilde{h}_{d}^{0}, \tilde{h}_{d}^{-} ight)$	$(1,2,rac{1}{2})$	
Gauge Supermultiplets of the MSSM					
		spin 1/2	spin 1	$\mathrm{SU}(3)_c, \mathrm{SU}(2)_L, \mathrm{U}(1)_Y$	
		$ ilde{g}$.	g	$({\bf 8},{f 1},0)$	
		$\tilde{W}^{\pm}, \tilde{W}^{0}$	W^{\pm}_{μ}, W^0_{μ}	$({f 1},{f 3},0)$	
		\tilde{B}	B_{μ}	(1, 1, 0)	

W. G. H. vacuos constraints

Electroweak symmetry breaking is automatic

- Higgs potential determined by theory, not put by hand
- Higgs mass calculated, not an input
- ullet potential stable up to high scales, quartics $\sim g_1^2+g_2^2$

There are, however, problems

- (tree level) Higgs mass bounded by M_Z
- need (large) loop corrections to shift $m_h^0 \rightarrow 125 \,\mathrm{GeV}$
- Supersymmetry apparently not exact in nature @ EW scale

Higgs potential of 2HDM type II

$$V = m_{11}^{2} H_{d}^{\dagger} H_{d} + m_{22}^{2} H_{u}^{\dagger} H_{u} + (m_{12}^{2} H_{u} \cdot H_{d} + h.c.) + \frac{\lambda_{1}}{2} (H_{d}^{\dagger} H_{d})^{2} + \frac{\lambda_{2}}{2} (H_{u}^{\dagger} H_{u})^{2} + \lambda_{3} (H_{u}^{\dagger} H_{u}) (H_{d}^{\dagger} H_{d}) + \lambda_{4} (H_{u}^{\dagger} H_{d}) (H_{d}^{\dagger} H_{u}) + \{\lambda_{5}, \lambda_{6}, \lambda_{7}\}$$

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$$m_{11}^2 = |\mu|^2 + m_{H_d}^2, \qquad \lambda_1 = \lambda_2 = -\lambda_3 = \frac{g_1^2 + g_2^2}{4},$$

$$m_{22}^2 = |\mu|^2 + m_{H_u}^2, \qquad \lambda_4 = \frac{g^2}{2},$$

$$m_{12}^2 = B_{\mu}, \qquad \lambda_5 = \lambda_6 = \lambda_7 = 0.$$

Desired = constructed

$$V = (m_{H_u}^2 + |\mu|^2)|H_u|^2 + (m_{H_d}^2 + |\mu|^2)|H_d|^2 - 2\operatorname{Re}(B_{\mu}H_u \cdot H_d) + \frac{g_1^2 + g_2^2}{8} \left(|H_u|^2 - |H_d|^2\right)^2 + \frac{g_2^2}{2}|H_d^{\dagger}H_u|^2$$

with

$$\langle H_u \rangle = \begin{pmatrix} 0 \\ v_u \end{pmatrix}, \langle H_d \rangle = \begin{pmatrix} v_d \\ 0 \end{pmatrix},$$

and
$$v_u^2 + v_d^2 = (174 \,\text{GeV})^2$$
, $v_u/v_d = \tan \beta$.

How to?

$$\frac{\partial V}{\partial h_u^0}\Big|_{\substack{h_{u,d} \to v_{u,d}}} = 2(\boldsymbol{m}_{\boldsymbol{H}_u}^2 + |\mu|^2)v_u - 2\operatorname{Re} B_{\mu}v_d + \frac{g_1^2 + g_2^2}{4}(v_u^2 - v_d^2)v_u$$
$$\frac{\partial V}{\partial h_d^0}\Big|_{\substack{h_{u,d} \to v_{u,d}}} = 2(\boldsymbol{m}_{\boldsymbol{H}_d}^2 + |\mu|^2)v_d - 2\operatorname{Re} B_{\mu}v_u - \frac{g_1^2 + g_2^2}{4}(v_u^2 - v_d^2)v_d$$

Many scalars from SUSY

The Minimal Supersymmetric Standard Model: A multi-scalar theory

$$V = V_F + V_D + V_{\text{soft}}$$
with
$$V_F = \sum_i \left| \frac{\partial W}{\partial \phi_i} \right|^2,$$

$$V_D = \frac{1}{2} \sum_a g_a^2 \left(\sum_i \phi_i^{\dagger} T^a \phi_i \right)^2$$

$$V_{\text{soft}} = \sum_i m_{\phi_i}^2 |\phi_i|^2 + \sum_{ijk} A_{ik}^{(j)} \phi_i^{\dagger} \phi_j \phi_k$$

 \leftrightarrow

The Standard Model: A single scalar theory

$$V_{\mathsf{SM}} = -\mu^2 H^\dagger H + rac{\lambda}{4} \left(H^\dagger H
ight)^2$$

W. G. H. vacuos constraints

The field content of the MSSM...

$$\mathcal{W} = \mu \ H_d \cdot H_u + Y_{ij}^u \ H_u \cdot Q_i \bar{T}_j - Y_{ij}^d \ H_d \cdot Q_i \bar{B}_j - Y_{ij}^\ell \ H_d \cdot L_i \bar{E}_j,$$

i, j = 1, ..., 3 generations; $H_d \cdot H_u = h_d^- h_u^+ - h_d^0 h_u^0$.

Soft SUSY breaking terms for the scalars...

$$\begin{split} \mathcal{L}_{\text{soft}} &= m_{H_d}^2 \, |h_d|^2 + m_{H_u}^2 \, |h_u|^2 \\ &+ \left(\tilde{m}_Q^2 \right)_{ij} \, \tilde{Q}_i^* \tilde{Q}_j + \left(\tilde{m}_u^2 \right)_{ij} \, \tilde{u}_i^* \tilde{u}_j + \left(\tilde{m}_d^2 \right)_{ij} \, \tilde{d}_i^* \tilde{d}_j \\ &+ \left(\tilde{m}_L^2 \right)_{ij} \, \tilde{L}_i^* \tilde{L}_j + \left(\tilde{m}_e^2 \right)_{ij} \, \tilde{e}_i^* \tilde{e}_j \\ &+ A_{ij}^u \, h_u \cdot \tilde{Q} \, \tilde{u}^* - A_{ij}^d \, h_d \cdot \tilde{Q} \, \tilde{d}^* \\ &- A_{ij}^\ell \, h_d \cdot \tilde{Q} \, \tilde{e}^* \end{split}$$

- about 100 new parameters
- in the scalar sector only!



[commons.wikimedia.org]

W. G. H. vacuos constraints

SOFT SUSY breaking complicates everything

A multi-scalar theory

- 2 Higgs doublets
- 2×6 scalar quarks, 6 + 3 scalar leptons
- 12 colored and 18 + 2 charged directions
- charged Higgs directions "safe"
- SM Higgs potential: SO(4) symmetry

The hazard

- impossible to minimize directly, analytically
- colored directions sensitive to all kinds of SUSY breaking
- spontaneous breaking of color charge: $\langle \tilde{q} \rangle \neq 0$

The true vacuum

- effective potential: average energy density
- global minimum: true ground state of the theory

[Casas et al. 1996]

The third generation MSSM

$$\mathcal{W} = \mu H_d \cdot H_u + y_t H_u \cdot Q\bar{T} - y_b H_d \cdot Q\bar{B}$$

- \bullet large couplings to Higgs doublets (y_t and y_b comparably large)
- large stop contribution (X_t, A_t) to light Higgs mass needed
- an eta resummation for m_b influences y_b

Properties of the (effective) scalar potential

- no UFB directions (due to quantum corrections)
- D-terms: (comparably) large contributions ϕ^4
- "dangerous" directions: small quadrilinears + large trilinears

Analytic constraints

- define certain directions in field space: great simplification
- e.g. D-terms absent: $|\tilde{Q}_L| = |\tilde{t}_R| = |h_2|$ (possibly miss sth.)

$$\begin{split} V_{\tilde{q},h} &= \tilde{t}_L^* \left(\tilde{m}_L^2 + |y_t h_u|^2 \right) \tilde{t}_L + \tilde{t}_R^* \left(\tilde{m}_t^2 + |y_t h_u|^2 \right) \tilde{t}_R \\ &+ \tilde{b}_L^* \left(\tilde{m}_L^2 + |y_b h_d|^2 \right) \tilde{b}_L + \tilde{b}_R^* \left(\tilde{m}_b^2 + |y_b h_d|^2 \right) \tilde{b}_R \\ &- \left[\tilde{t}_L^* \left(\mu^* y_t \ h_d^* - A_t h_u \right) \tilde{t}_R + \text{h.c.} \right] \\ &- \left[\tilde{b}_L^* \left(\mu^* y_b \ h_u^* - A_b h_d \right) \tilde{b}_R + \text{h.c.} \right] \\ &+ |y_t|^2 |\tilde{t}_L|^2 |\tilde{t}_R|^2 + |y_b|^2 |\tilde{b}_L|^2 |\tilde{b}_R|^2 \\ &+ \frac{g_1^2}{8} \left(|h_u|^2 - |h_d|^2 + \frac{1}{3} |\tilde{b}_L|^2 + \frac{2}{3} |\tilde{b}_R|^2 + \frac{1}{3} |\tilde{t}_L|^2 - \frac{4}{3} |\tilde{t}_R|^2 \right)^2 \\ &+ \frac{g_2^2}{8} \left(|h_u|^2 - |h_d|^2 + |\tilde{b}_L|^2 - |\tilde{t}_L|^2 \right)^2 \\ &+ \frac{g_3^2}{8} \left(|\tilde{t}_L|^2 - |\tilde{t}_R|^2 + |\tilde{b}_L|^2 - |\tilde{b}_R|^2 \right)^2 \\ &+ (m_{h_u}^2 + |\mu|^2) |h_u|^2 + (m_{h_d}^2 + |\mu|^2) |h_d|^2 - 2 \operatorname{Re}(B_\mu \ h_d h_u). \end{split}$$

$$\begin{split} V_{\tilde{q},h} &= \tilde{t}_L^* \left(\tilde{m}_L^2 + |y_t h_u|^2 \right) \tilde{t}_L + \tilde{t}_R^* \left(\tilde{m}_t^2 + |y_t h_u|^2 \right) \tilde{t}_R \\ &+ \tilde{b}_L^* \left(\tilde{m}_L^2 + |y_b h_d|^2 \right) \tilde{b}_L + \tilde{b}_R^* \left(\tilde{m}_b^2 + |y_b h_d|^2 \right) \tilde{b}_R \\ &- \left[\tilde{t}_L^* \left(\mu^* y_t \ h_d^* - A_t h_u \right) \tilde{t}_R + \text{h.c.} \right] \\ &- \left[\tilde{b}_L^* \left(\mu^* y_b \ h_u^* - A_b h_d \right) \tilde{b}_R + \text{h.c.} \right] \\ &+ |y_t|^2 |\tilde{t}_L|^2 |\tilde{t}_R|^2 + |y_b|^2 |\tilde{b}_L|^2 |\tilde{b}_R|^2 \\ &+ \frac{g_1^2}{8} \left(|h_u|^2 - |h_d|^2 + \frac{1}{3} |\tilde{b}_L|^2 + \frac{2}{3} |\tilde{b}_R|^2 + \frac{1}{3} |\tilde{t}_L|^2 - \frac{4}{3} |\tilde{t}_R|^2 \right)^2 \\ &+ \frac{g_2^2}{8} \left(|h_u|^2 - |h_d|^2 + |\tilde{b}_L|^2 - |\tilde{t}_L|^2 \right)^2 \\ &+ \frac{g_3^2}{8} \left(|\tilde{t}_L|^2 - |\tilde{t}_R|^2 + |\tilde{b}_L|^2 - |\tilde{b}_R|^2 \right)^2 \\ &+ (m_{h_u}^2 + |\mu|^2) |h_u|^2 + (m_{h_d}^2 + |\mu|^2) |h_d|^2 - 2 \operatorname{Re}(B_\mu \ h_d h_u). \end{split}$$

 $|\tilde{t}_L| = |\tilde{t}_R| = |\tilde{t}|, \ |\tilde{b}_L| = |\tilde{b}_R| = |\tilde{b}|$

$$\begin{split} V_{\tilde{q},h} &= \tilde{t}^{*} \left(\tilde{m}_{L}^{2} + |y_{t}h_{u}|^{2} \right) \tilde{t}^{*} + \tilde{t}^{*} \left(\tilde{m}_{t}^{2} + |y_{t}h_{u}|^{2} \right) \tilde{t} \\ &+ \tilde{b}^{*} \left(\tilde{m}_{L}^{2} + |y_{b}h_{d}|^{2} \right) \tilde{b}^{*} + \tilde{b}^{*} \left(\tilde{m}_{b}^{2} + |y_{b}h_{d}|^{2} \right) \tilde{b} \\ &- \left[\tilde{t}^{*} \left(\mu^{*}y_{t} h_{d}^{*} - A_{t}h_{u} \right) \tilde{t}^{*} + \text{h.c.} \right] \\ &- \left[\tilde{b}^{*} \left(\mu^{*}y_{b} h_{u}^{*} - A_{b}h_{d} \right) \tilde{b}^{*} + \text{h.c.} \right] \\ &+ |y_{t}|^{2} |\tilde{t}^{*}|^{2} |\tilde{t}^{*}|^{2} + |y_{b}|^{2} |\tilde{b}^{*}|^{2} \tilde{b}^{*}|^{2} \\ &+ \frac{g_{1}^{2}}{8} \left(|h_{u}|^{2} - |h_{d}|^{2} + |\tilde{b}^{*}|^{2} - |\tilde{t}^{*}|^{2} \right)^{2} \end{split}$$

 $+\,(m_{h_u}^2+|\mu|^2)|h_u|^2+(m_{h_d}^2+|\mu|^2)|h_d|^2-2\,{\rm Re}(B_\mu\;h_dh_u).$

 $|\tilde{t}_L| = |\tilde{t}_R| = |\tilde{t}|, \ |\tilde{b}_L| = |\tilde{b}_R| = |\tilde{b}|$

$$\begin{split} V_{\tilde{q},h} &= \tilde{t}^{*} \left(\tilde{m}_{L}^{2} + |y_{t}h_{u}|^{2} \right) \tilde{t}^{*} + \tilde{t}^{*} \left(\tilde{m}_{t}^{2} + |y_{t}h_{u}|^{2} \right) \tilde{t} \\ &+ \tilde{b}^{*} \left(\tilde{m}_{L}^{2} + |y_{b}h_{d}|^{2} \right) \tilde{b}^{*} + \tilde{b}^{*} \left(\tilde{m}_{b}^{2} + |y_{b}h_{d}|^{2} \right) \tilde{b} \\ &- \left[\tilde{t}^{*} \left(\mu^{*}y_{t} h_{d}^{*} - A_{t}h_{u} \right) \tilde{t}^{*} + \text{h.c.} \right] \\ &- \left[\tilde{b}^{*} \left(\mu^{*}y_{b} h_{u}^{*} - A_{b}h_{d} \right) \tilde{b}^{*} + \text{h.c.} \right] \\ &+ |y_{t}|^{2} |\tilde{t}^{*}|^{2} |\tilde{t}^{*}|^{2} + |y_{b}|^{2} |\tilde{b}^{*}|^{2} \tilde{b}^{*}|^{2} \\ &+ \frac{g_{1}^{2}}{8} \left(|h_{u}|^{2} - |h_{d}|^{2} + |\tilde{b}^{*}|^{2} - |\tilde{t}^{*}|^{2} \right)^{2} \end{split}$$

 $+ (m_{h_u}^2 + |\mu|^2)|h_u|^2 + (m_{h_d}^2 + |\mu|^2)|h_d|^2 - 2\operatorname{Re}(B_\mu h_d h_u).$

 $|\tilde{t}_L| = |\tilde{t}_R| = |\tilde{t}|, \ |\tilde{b}_L| = |\tilde{b}_R| = |\tilde{b}|; \ |\tilde{b}| = |h_d| = |\phi_1|, \ |\tilde{t}| = |h_u| = |\phi_2|$ W. G. H. vacuus constraints

$$\begin{split} V_{\tilde{q},h} &= \phi_2^* \left(\tilde{m}_L^2 + |y_t \phi_2|^2 \right) \phi_2 + \phi_2^* \left(\tilde{m}_t^2 + |y_t \phi_2|^2 \right) \phi_2 \\ &+ \phi_1^* \left(\tilde{m}_L^2 + |y_b \phi_1|^2 \right) \phi_1 + \phi_1^* \left(\tilde{m}_b^2 + |y_b \phi_1|^2 \right) \phi_1 \\ &- \left[\phi_2^* \left(\mu^* y_t \ \phi_1^* - A_t \phi_2 \right) \phi_2 + \text{h.c.} \right] \\ &- \left[\phi_1^* \left(\mu^* y_b \ \phi_2^* - A_b \phi_1 \right) \phi_1 + \text{h.c.} \right] \\ &+ |y_t|^2 |\phi_2|^2 |\phi_2|^2 + |y_b|^2 |\phi_1|^2 |\phi_1|^2 \end{split}$$

$$+ (m_{h_u}^2 + |\mu|^2) |\phi_2|^2 + (m_{h_d}^2 + |\mu|^2) |\phi_1|^2 - 2 \operatorname{Re}(B_\mu \ \phi_1 \phi_2).$$

$$\tilde{t}_L| = |\tilde{t}_R| = |\tilde{t}|, \ |\tilde{b}_L| = |\tilde{b}_R| = |\tilde{b}| \ ; \ |\tilde{b}| = |h_d| = |\phi_1|, \ |\tilde{t}| = |h_u| = |\phi_2|$$

W. G. H. vacuos constraints

$$V_{\tilde{q},h} = \phi_2^* \left(\tilde{m}_L^2 + |y_t \phi_2|^2 \right) \phi_2 + \phi_2^* \left(\tilde{m}_t^2 + |y_t \phi_2|^2 \right) \phi_2$$
$$- \left[\phi_2^* \left(-A_t \phi_2 \right) \phi_2 + \text{h.c.} \right]$$
$$+ |y_t|^2 |\phi_2|^2 |\phi_2|^2$$

$$+ (m_{h_{u}}^{2} + |\mu|^{2})|\phi_{2}|^{2}$$

$$\tilde{t}_{L}| = |\tilde{t}_{R}| = |\tilde{t}|, |\tilde{b}_{L}| = |\tilde{b}_{R}| = |\tilde{b}|; \quad b = h_{d} = \phi_{1}, \quad |\tilde{t}| = |h_{u}| = |\phi_{2}|$$
W. G. H. values constraints

Minimize the potential

$$V(\phi) = m^2 \phi^2 - A \phi^3 + \lambda \phi^4,$$

with $m^2=m_{h_2}^2+|\mu|^2+\tilde{m}_L^2+\tilde{m}_t^2,\,A=-A_t$ and $\lambda=3y_t^2.$

Mathematics for the kindergarden

Minimize the potential

$$V(\phi) = m^2 \phi^2 - A \phi^3 + \lambda \phi^4,$$

with $m^2 = m_{h_2}^2 + |\mu|^2 + \tilde{m}_L^2 + \tilde{m}_t^2$, $A = -A_t$ and $\lambda = 3y_t^2$.
Answer:

$$\phi_0 = 0, \qquad \phi_{\pm} = \frac{3A \pm \sqrt{9A^2 32\lambda m^2}}{8\lambda}.$$

Condition to be safe from non-standard (i.e. non-trivial) minima:

$$V(\phi_{\pm}) > 0 \quad \hookrightarrow \quad m^2 > \frac{A^2}{4\lambda}$$

Mathematics for the kindergarden

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$$V(\phi_{\pm}) > 0 \quad \hookrightarrow \quad m^2 > \frac{A^2}{4\lambda}$$

Well-known constraints

[Gunion, Haber, Sher '88]

$$\begin{split} |A_t|^2 < 3y_t^2 \left(m_{h_2}^2 + |\mu|^2 + \tilde{m}_L^2 + \tilde{m}_t^2\right) \\ |A_b|^2 < 3y_b^2 \left(m_{h_1}^2 + |\mu|^2 + \tilde{m}_L^2 + \tilde{m}_b^2\right) \end{split}$$
 for the limiting cases $|\tilde{t}_L| = |\tilde{t}_R| = |h_2|$ and $|\tilde{b}_L| = |\tilde{b}_R| = |h_1|!$







Problem already known for a while

- problem noticed
- "A-parameter bounds"

- [Frere, Jones, Raby '83]
- [Gunion, Haber, Sher '87]
- classification of all dangerous directions [Casas, Lleyda, Muñoz '96]
- including flavor violation

[Casas and Dimopuolos '96]

Stability \neq no Instability \Rightarrow Metastability

Vacuum tunneling

[Kusenko, Langacker '96; Blinov, Morissey '13]

The tool

VeVacious

[Camargo-Molina, O'Leary, Porod, Staub '13]

- finds all (?) tree-level minima
- minimizes scalar potential in the vicinity at one loop
- calculates bounce action / tunneling times [CosmoTransitions]

- We have discovered the Higgs!
- No sign of SUSY so far...
- $\bigcirc \ m_h = 125 \, {\rm GeV}$

(all SUSY literature during LEP era expected it to be $\lesssim 100\,{\rm GeV})$

- Onsequently: large radiative corrections!
- large stop mixing needed? heavy SUSY spectrum? (or hidden in some hardly accessible valley)
- approach today:
 - less focused on unified models
 - still certain scenarios
 - $\tan\beta$ resummation for bottom quark mass (large $\tan\beta$)
 - low $\tan\beta$ favored (for $M_A \lesssim 800 \, {
 m GeV}$, direct search $A \to \tau \tau$)

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Semi-analytical bounds/exclusions important for fast processing!

My "pMSSM"

- no unification (more than m_0 , $m_{1/2}$, A_0 , aneta and ${
 m sign}\,\mu$)
- free parameters: (although similar choice as in CMSSM)

•
$$\tilde{m}_L^2 = \tilde{m}_t^2 = \tilde{m}_b^2 = M_{\text{SUSY}}^2$$

- μ , $\tan\beta$
- A_t , A_b (not necessarily equal)
- $m^2_{h_{1,2}}$ determined from ew breaking, B_μ related to M_A
- no RG running needed = parameters taken at the SUSY scale

Why no RG-improvement?

- SUSY scale parameters; limits on this parameters
- destabilization of ew vacuum around SUSY scale
- no Planck scale vevs! (maybe there are...in addition)
- in the spirit of the pMSSM as phenomenologically as possible
- only small RG modifications, qualitative features unchanged

More freedom!

Less constraints, more parameters, more fields, more vevs...

Restriction to certain directions too restrictive!

- give up $|\tilde{q}_L| = |\tilde{t}_R| = |h_2|$
- allow for $h_1 \neq 0$ and $\tilde{b} \neq 0$
- "break" $\tilde{q}_L \rightarrow \tilde{t}_L + \tilde{b}_L$
- back to full scalar potential!

Simplify your life

- $h_2 = \phi$
- $|\tilde{t}_L| = |\tilde{t}_R| = |\tilde{t}| = \alpha |\phi|$
- $h_1 = \eta \phi$
- $|\tilde{b}_L| = |\tilde{b}_R| = |\tilde{b}| = \beta |\phi|$
- all fields and parameters real, $\alpha,\beta>0,\,\eta\in\mathbb{R}$
- SU(3)_c-flatness: $\tilde{t}_L = \tilde{t}_R$ and $\tilde{b}_L = \tilde{b}_R$

A simple view of a complicated object

$$\begin{split} h_{2} &= \phi, \qquad |\tilde{t}| = \alpha |\phi|, \qquad h_{1} = \eta \phi, \qquad |\tilde{b}| = \beta |\phi| \\ V_{\phi} &= \left(m_{h_{2}}^{2} + \eta^{2} m_{h_{1}}^{2} + (1 + \eta^{2}) \mu^{2} - 2B_{\mu} \eta \right. \\ &+ \left(\alpha^{2} + \beta^{2}\right) \tilde{m}_{L}^{2} + \alpha^{2} \tilde{m}_{t}^{2} + \beta^{2} \tilde{m}_{b}^{2}\right) \phi^{2} \\ &- 2 \left(\alpha^{2} (\mu y_{t} \eta - A_{t}) + \beta^{2} (\mu y_{t} - \eta A_{b})\right) \phi^{3} + \left(\alpha^{2} y_{t}^{2} + \beta^{4} y_{b}^{2}\right) \phi^{4} \\ &+ \left(\frac{g_{1}^{2} + g_{2}^{2}}{8} (1 - \eta^{2} + \beta^{2} - \alpha^{2})^{2} + 2\alpha^{2} y_{t}^{2} + 2\beta^{2} y_{b}^{2}\right) \phi^{4} \\ &\equiv M^{2} (\eta, \alpha, \beta) \phi^{2} - \mathcal{A} (\eta, \alpha, \beta) \phi^{3} + \lambda (\eta, \alpha, \beta) \phi^{4}, \end{split}$$

with

$$\begin{split} M^2 &= m_{h_2}^2 + \eta^2 m_{h_1}^2 + (1+\eta^2) \mu^2 - 2B_\mu \eta \\ &+ (\alpha^2 + \beta^2) \tilde{m}_L^2 + \alpha^2 \tilde{m}_t^2 + \beta^2 \tilde{m}_b^2 \,, \end{split}$$

$$\mathcal{A} = 2\alpha^2 \eta \mu y_t - 2\alpha^2 A_t + 2\beta^2 \mu y_b - 2\eta \beta^2 A_b \,,$$

$$\begin{split} \lambda &= \frac{g_1^2 + g_2^2}{8} (1 - \eta^2 + \beta^2 - \alpha^2)^2 \\ &+ (2 + \alpha^2) \alpha^2 y_t^2 + (2\eta^2 + \beta^2) \beta^2 y_b^2 \,. \end{split}$$

Optimized Charge and Color Breaking

[Gunion, Haber, Sher '88; Casas, Lleyda, Muñoz '96]

$$\begin{aligned} \text{The same but different ("A-parameter bounds")}} \\ & \mathcal{A}^2 < 4\lambda M^2 \\ & \downarrow \\ 4\min_{\{\eta,\alpha,\beta\}} \lambda(\eta,\alpha,\beta) M^2(\eta,\alpha,\beta) > \max_{\{\eta,\alpha,\beta\}} (\mathcal{A}(\eta,\alpha,\beta))^2 \\ \hline & \boldsymbol{h_u} = \tilde{b}, \, \boldsymbol{h_d}^0 = 0 & \text{[WGH'15]} \\ & \boldsymbol{m_{H_u}^2} + \mu^2 + \tilde{m}_Q^2 + \tilde{m}_b^2 > \frac{(\mu y_b)^2}{y_b^2 + (g_1^2 + g_2^2)/2} \\ \hline & \textbf{|h_d|^2} = |\boldsymbol{h_u}|^2 + |\tilde{b}|^2, \, \tilde{b} = \alpha h_u & \text{[WGH'15]} \\ & \boldsymbol{m_{11}^2(1 + \alpha^2)} + m_{22}^2 \pm 2m_{12}^2\sqrt{1 + \alpha^2} + \alpha^2(\tilde{m}_Q^2 + \tilde{m}_b^2) > \frac{4\mu^2\alpha^2}{2 + 3\alpha^2} \end{aligned}$$

W. G. H. vacuos constraints

The issue of including field directions $ilde{t} = 0, h_d = 0, A_b = 0$



W. G. H. vacuos constraints
The issue of including field directions



 $ilde{t}=0, A_b=0$

The issue of including field directions



W. G. H. vacuos constraints

 $A_b = 0$

The issue of including field directions



 $A_b = A_t$

Higgs mass bounded by M_Z

$$m_{h^0} \le M_Z^2 \cos^2(2\beta)$$

large loop contribution in the MSSM needed

$$\Delta m_{h^0}^2 = \frac{3v^2}{4\pi^2} y_t^4 \left[\log\left(\frac{M_{\mathsf{SUSY}}^2}{m_t^2}\right) + \frac{X_t^2}{M_{\mathsf{SUSY}}^2} \left(1 - \frac{X_t^2}{12M_{\mathsf{SUSY}}^2}\right) \right]$$

• dominant contribution with y_t , $X_t = (A_t/y_t - \mu \cot \beta)$

• large A_t and/or M_{SUSY} needed

Section Pres

Closing in on the parameter space









Open up parametrs space increasing MSUSY



Open up parametrs space increasing MSUSY



W. G. H. vacuos constraints

 $A_b = 0$



W. G. H. vacuos constraints

 $\mu = 350\,{
m GeV}$



W. G. H. vacuos constraints

 $\mu = M_{\rm MSUSY}$

 $\mu = 500\,{
m GeV}$



 X_t/M_{SUSY}



W. G. H. vacuos constraints

M_{SUSY}/GeV

$\mu = -500 \,\mathrm{GeV}$

Why next?

- Why not?
- Add a SM singlet superfield.
- Richer phenomenology (Higgs and neutralino sector)

The NMSSM solves the " μ -problem"

$$\mathcal{W}_{\mathsf{MSSM}} = \mu \; H_u \cdot H_d + \mathsf{Yukawa}$$

only dimensionful parameter μ has to be \sim electroweak scale

$$\mathcal{W}_{\text{NMSSM}} \supset \lambda \ SH_u \cdot H_d + \frac{\kappa}{3}S^3$$

dynamical $\mu\text{-term:}\;\lambda\langle S\rangle=\mu_{\rm eff}$

 \mathbb{Z}_3 symmetry forbids dimensionful couplings (bilinear, tadpole terms)

Alternative vevs in the NMSSM

Trilinear terms in the Higgs sector

$$\begin{split} V_{\rm soft} &= m_{H_u}^2 |h_u|^2 + m_{H_d}^2 |h_d|^2 + m_S^2 |s|^2 \\ &+ \left(A_\lambda \lambda \; sh_u \cdot h_d + \frac{1}{3} A_\kappa \kappa \; s^3 + {\rm h.\, c.} \; \right) \end{split}$$

 V_{Higgs} is a *multivariate* polynomial to order 4

$$V_{\mathsf{Higgs}} \subset \{h_u^2, h_d^2, s^2, sh_uh_d, s^2, s^2h_uh_d, s^2h_u^2, s^2h_d^2, h_u^2h_d^2, h_u^4, h_d^4, s^4\}$$

Not only one minimum! V_{ew} not necessarily the global minimum!

Minimisation conditions are in general misleading!

$$\frac{\partial V}{\partial h_u}\Big|_{\mathsf{vev}} = 2m_{H_u}^2 v_u + \dots$$
$$\frac{\partial V}{\partial h_d}\Big|_{\mathsf{vev}} = 2m_{H_d}^2 v_d + \dots$$
$$\frac{\partial V}{\partial h_u}\Big|_{\mathsf{vev}} = 2m_S^2 v_s + \dots$$

linear equations for soft SUSY breaking masses $m^2_{H_u}$, $m^2_{H_d}$, m^2_S , can be solved uniquely; determine numerical values for those

The crazy world of the NMSSM

Simplification: h_u^0, h_d^0, s^0 only (three fields, many vacua), real fields and parameters

$$\begin{split} V_{\text{Higgs}} &= \boldsymbol{m_{H_u}^2} \ h_u^2 + \boldsymbol{m_{H_d}^2} \ h_d^2 + \boldsymbol{m_S^2} \ s^2 \\ &+ \frac{2}{3} \kappa A_{\kappa} \ s^3 + 2\lambda A_{\lambda} \ sh_u h_d \\ &+ \left(\kappa \ s^2 - \lambda \ h_u h_d\right)^2 + \lambda^2 \ s^2 \left(h_u^2 + h_d^2\right) \\ &+ \frac{g_1^2 + g_2^2}{8} \left(h_u^2 - h_d^2\right)^2 \end{split}$$

However

Solutions for minimization equations with $\langle h_u \rangle \neq v_u$, $\langle h_d \rangle \neq v_d$ and $\langle s \rangle \neq \mu_{\rm eff} / \lambda$ possible, viable, existing *and* leading to a true vacuum.

Potential value at the minimum to be compared with

$$V_{\min}^{\text{des}} = -\frac{g_1^2 + g_2^2}{8} v^4 \cos^2(2\beta) - \frac{\lambda^2}{4} v^4 \sin^2(2\beta) - \frac{\kappa^2}{\lambda^4} \mu_{\text{eff}}^4 - v^2 \mu_{\text{eff}}^2 \left(1 - \frac{\kappa^2}{\lambda^2} \sin(2\beta)\right) - \frac{1}{3} \frac{\kappa A_\kappa}{\lambda^3} \mu_{\text{eff}}^3 + \frac{v^2}{2} A_\lambda \mu_{\text{eff}} \sin(2\beta)$$

Constraint on A_{κ}

$$A_{\kappa}^2 > 9m_S^2$$

 $A_\kappa^2 < 8m_S^2$: no $\langle s
angle
eq 0$. [Derendinger, Savoy '84; Ellwanger et al. '97]

"Tachyonic" Higgs masses

- "problem" of tachyonic masses well known
- \bullet one mass eigenvalue of \mathcal{M}^2_S , \mathcal{M}^2_P or charged Higgs mass $m^2_{H^\pm}$ negative
- tachyonic mass = negative curvature = alternative vev (!)

However

Careful analysis shows that tachyonic masses are not enough! [see e.g. Kanehata, Kobayashi, Konishi, Seto, Shimomura '11]



















Conclusions

The SUSY paradigm and a $125\,{\rm GeV}$ Higgs. . .

- leads to severe constraints in MSSM parameterspace
- theoretical consistency: Higgs vacuum is true vacuum
- vacuum tunneling neglected/ignored

Extending the MSSM

- Next-to-Minimal: add gauge singlet
- changes phenomenology dramatically
- tachyonic Higgses and/or alternative Higgs vevs

The Cosmic Higgs connection

[CHICO, canceled :(]

- possibility of incproporating Higgs inflation in the (N)MSSM
- Singlet stabilizes inflationary trajectory
- NMSSM + $\mu H_d \cdot H_u$, changes phenomenology drastically

local $U(1) \mathcal{R}$ symmetry

- \bullet part of the superconformal ${\rm SU}(2,2|1)$
- χ term breaks continous ${\mathcal R}$ and discrete ${\mathbb Z}_3$ symmetry
- breaking at dimension 6: $\sim \chi \frac{\lambda^2 h^6}{M_P^2}$

Corrected Superpotential

$$\begin{split} \mathcal{W}_{\text{eff}} &\to \mathcal{W}e^{X(\Phi)/M_P^2} = \mathcal{W} + \frac{\langle \mathcal{W}_{\text{hid}} \rangle}{M_P^2} X(\Phi) \\ &\simeq \mathcal{W} + m_{3/2} X(\Phi) \end{split}$$

The iNMSSM

$$\mathcal{W}_{\mathsf{eff}} = \lambda \ SH_u \cdot H_d + \frac{\kappa}{3} \ S^3 + \frac{3}{2} \chi m_{3/2} H_u \cdot H_d$$

Cosmo pheno requires $\chi \simeq 10^5 \lambda$

Phenomenology of the inflationary term

like the NMSSM with an extended effective μ term

$$\mu_{\rm eff}' = \lambda \langle S \rangle + \frac{3}{2} \chi m_{3/2} = \mu_{\rm eff} + \mu$$

Additional soft SUSY breaking term

$$\begin{split} V_{\mathsf{soft}} = & \lambda A_{\lambda} \ SH_u \cdot H_d + \frac{1}{3} \kappa A_{\kappa} \ S^3 \\ & + \frac{3}{2} B_{\mu} \chi m_{3/2} \left(H_u \cdot H_d + \mathsf{h.c.} \right) \end{split}$$

Higgs potential of the iNMSSM

$$\begin{split} V &= \left[m_{H_d}^2 + \left(\mu + \lambda S \right)^2 \right] |H_d|^2 + \left[m_{H_u}^2 + \left(\mu + \lambda S \right)^2 \right] |H_u|^2 + m_S^2 S^2 \\ &+ \frac{2}{3} \kappa A_\kappa S^3 + \left[\kappa S^2 + \lambda H_u \cdot H_d \right]^2 + 2 \left(B_\mu \mu + \lambda A_\lambda S \right) H_u \cdot H_d \\ &+ \frac{g_1^2 + g_2^2}{8} \left(|H_d|^2 - |H_u|^2 \right)^2 + \frac{g_2^2}{2} |H_d^{\dagger} H_u|^2 \end{split}$$









- \bullet (heavy) Higgs decays depending on μ and $\mu+\mu_{\rm eff}$
- Neutralino sector:
 - $\bullet \ \ {\rm Higgsinos} \sim (\mu + \mu_{\rm eff})$
 - Singlino $\sim \frac{\kappa}{\lambda} \mu_{\rm eff}$
- However, μ can also be small

[Ferrara, Kallosh, Linde, Marrani, Van Proeyen 2010]

• Work in progress ...

- A new view on vacuum stability in the MSSM, JHEP 08 (2016) 126, Wolfgang G. Hollik [1606.08356]
- [2] Charge and color breaking constraints in the Minimal Supersymmetric Standard Model associated with the bottom Yukawa coupling, Physics Letters B 752 7–12 (2016), Wolfgang G. Hollik [1508.07201]
- [3] Vacuum stability of the effective Higgs potential in the Minimal Supersymmetric Standard Model, Physical Review D 90 035025 (2014) Markus Bobrowski, Guillaume Chalons, Wolfgang G. Hollik, Ulrich Nierste [1407.2814]

Work in progress...

- Higgs Phenomenology of NMSSM inflation, with Stefan Liebler, Gudrid Moortgat-Pick, Sebastian Paßehr, Georg Weiglein, arXiv:1711.0xxxx
- [2] The Higgs vacuum in the NMSSM, with Georg Weiglein, Jonas Wittbrod arXiv:171x.0yyyy
Backup

Slides

W. G. H. vacuos constraints

$\tan\beta$ resummation for bottom yukawa coupling

Yukawa coupling not directly proportional to mass (same for y_t) $y_b = \frac{m_b}{v_d(1+\Delta_b)}$

[Hall, Rattazzi, Sarid '94; Carena, Garcia, Nierste, Wagner '99]

$$\begin{split} \Delta_b^{\text{gluino}} &= \frac{2\alpha_s}{3\pi} \mu M_{\tilde{G}} \tan \beta C_0(\tilde{m}_{\tilde{b}_1}, \tilde{m}_{\tilde{b}_2}, M_{\tilde{G}}), \\ \Delta_b^{\text{higgsino}} &= \frac{Y_t^2}{16\pi^2} \mu A_t \tan \beta C_0(\tilde{m}_{\tilde{t}_1}, \tilde{m}_{\tilde{t}_2}, \mu). \end{split}$$



W. G. H. vacuos constraints