



Neutrino Masses & Mixings 2012

Guido Altarelli Universita' di Roma Tre CERN

In the last 2 decades data on ν oscillations have added some (badly needed) fresh experimental input to particle physics



reactors



atmosphere



accelerators

Sector 1

Schwetz

Homestake, SAGE, GALLEX SuperK, SNO, Borexino

KamLAND, CHOOZ

SuperKamiokande

K2K, MINOS, T2K

Frend

v masses are not all vanishing but they are very small

This suggests that ν 's are Majorana particles and that the lepton number L is not conserved



v mixing angles follow a different pattern from quark mixings

This also is probably related to the Majorana nature of v's



 $P(v_e < v_\mu) = |< v_\mu(L)| v_e > |^2 = sin^2(2\theta) \cdot sin^2(\Delta m^2 L/4E)$

At a distance L, v_{μ} from μ^{-} decay can produce e⁻ via charged weak interact's



Evidence for solar and atmosph. v oscillations confirmed on earth by K2K, KamLAND, MINOS, T2K...

 Δm^2 values: $\Delta m_{atm}^2 \sim 2.5 \ 10^{-3} \ eV^2$, $\Delta m_{sol}^2 \sim 8 \ 10^{-5} \ eV^2$ and mixing angles measur'd: θ_{12} (solar) large θ_{23} (atm) large~ maximal θ_{13} (T2K, MINOS, DOUBLE CHOOZ) small A 3rd frequency? A persisting confusion: LSND+MiniBooNE

Sterile (no weak int's) neutrino



Are sterile v's coming back? A number of "hints" (they do not make an evidence but pose an experimental problem that needs clarification)

- LSND and MiniBoone
- Reactor flux & anomaly
- Gallium v_e disappearance vs v_e^{bar} reactor limits

If all true (unlikely) then need at least 2 sterile $\nu ^{\prime }s$

Important information also from

Neutrino counting from cosmology





The reactor anomaly



Systematic errors not shown in this figure (estimated in paper)! Certainly of the same order of the shift. They could well be larger than estimated



This is the compromise realized in the fit



$$\chi^{2}_{min} = 59.8$$

NdF = 65
GoF = 66%
 $\sin^{2} 2\vartheta = 0.17$
 $\Delta m^{2} = 4.17 \text{ eV}^{2}$
PGoF = 1.1%

Cosmology could accept one sterile neutrino

The bound from nucleosynthesis is the most stringent (assuming thermal properties at decoupling)

► BBN: $N_s = 0.22 \pm 0.59$ [Cyburt, Fields, Olive, Skillman, AP 23 (2005) 313, astro-ph/0408033] $N_s = 0.64^{+0.40}_{-0.35}$ [Izotov, Thuan, ApJL 710 (2010) L67, arXiv:1001.4440]

▶ BBN: $N_s < 1.2 (95\% \text{ CL})$ Mangano, Serpico, 1103.1261

▶ BBN: N_s < 1.54 (95% CL) [M. Pettini, et al, arXiv:0805.0594]







In any case only a small leakage from active to sterile neutrinos is allowed by present data



Most common EW scale BSM do not contain sterile neutrinos. A sterile neutrino would probably be a remnant of some hidden sector or of gravity. So would be a great discovery.



Recent Fits (2011)

Quantity	Fogli et al ¹⁾	Schwetz et al ²⁾
$\Delta m_{sun}^2 \ (10^{-5} \ {\rm eV}^2)$	$7.58^{+0.22}_{-0.26}$	$7.59^{+0.20}_{-0.18}$
$\Delta m_{atm}^2 \ (10^{-3} \ {\rm eV}^2)$	$2.35^{+0.12}_{-0.09}$	$2.50^{+0.09}_{-0.16}$
$\sin^2 \theta_{12}$	$0.312^{+0.017}_{-0.016}$	$0.312^{+0.017}_{-0.015}$
$\sin^2 \theta_{23}$	$0.42^{+0.08}_{-0.03}$	$0.52^{+0.06}_{-0.07}$
$\sin^2 \theta_{13}$	0.025 ± 0.007	$0.013\substack{+0.007\\-0.005}$





Recent results on θ_{13} (T2K, MINOS, DOOBLE CHOOZ)



 \oplus

The near future of θ_{13}



v oscillations measure Δm^2 . What is m^2 ?

 $\Delta m_{atm}^2 \sim 2.5 \ 10^{-3} \ eV^2 = (0.05 \ eV)^2$; $\Delta m_{sun}^2 \sim 8 \ 10^{-5} \ eV^2 = (0.009 \ eV)^2$ End-point tritium Direct limits $m_{"ve"} < 2.2 \text{ eV}$ β decay (Mainz, Troitsk) Future: Katrin, MARE $m_{"vu"} < 170 \text{ KeV}$ 0.2 eV sensitivity $m_{ee} = |\sum U_{ei}^2 m_i|$ $m_{\nu \tau} < 18.2 MeV$ (Karsruhe) • 0νββ $m_{ee} < 0.2 - 0.7 - ? eV$ (nucl. matrix elmnts) Evidence of signal? **Klapdor-Kleingrothaus** $(h^2 \sim 1/2)$ Cosmology $\Omega_v h^2 \sim \Sigma_i m_i / 94 eV$ WMAP, SDSS, $\Sigma_i m_i < 0.2-0.7 \text{ eV}$ (dep. on data&priors) 2dFGRS, Ly- α Any v mass < 0.06 - 0.23 - 2.2 eV</p>

Current constraints on neutrino mass from Cosmology

By itself CMB (eg WMAP) is only mildly sensitive to $\Sigma_i m_i$ Only with Large Scale Structure the limit becomes stronger.



Dark Matter

WMAP, BAO....

Most of the Universe is not made up of atoms: $\Omega_{tot} \sim 1$, $\Omega_{b} \sim 0.045$, $\Omega_{m} \sim 0.27$ Most is Dark Matter and Dark Energy

Most Dark Matter is Cold (non relativistic at freeze out) Significant Hot Dark matter is disfavoured Hot Dark Matter does not "stick" enough at short distances (Galaxy haloes...)

 $\stackrel{\checkmark}{\sim}$ Neutrinos are not much cosmo-relevant: $\Omega_v < 0.015$





Neutrino masses are really special!

 $h_{\rm t}/(\Delta m_{\rm atm}^2)^{1/2} \sim 10^{12}$

Massless v's?

- no v_R
- L conserved

Small v masses?

- v_R very heavy
- L not conserved

Very likely: v's are special as they are Majorana fermions

Are neutrinos Dirac or Majorana fermions?

Under charge conjugation C: particle <--> antiparticle

For bosons there are many cases of particles that coincide (up to a phase) with their antiparticle:

 $\pi^0, \rho^0, \omega, \gamma, Z^0....$

A fermion that coincides with its antiparticle is called a Majorana fermion

Are there Majorana fermions? Neutrinos are probably Majorana fermions



uuuv
e
dddecccv
 μ
sss
 μ *tttv*
 τ
bbb

- Of all fundamental fermions only v's are neutral If lepton number L conservation is violated then no conserved charge distinguishes neutrinos from antineutrinos
 Majorana v's : neutrinos and antineutrinos coincide neutrinos are their own antiparticles
 - v's have very small masses The two facts are probably related



The field of an electron (massive, charged) has 4 components

In fact there are 4 dof: e^{-} , e^{+} , h = +, -(h is the helicity: component of spin along momentum)





A 2-component description is possible in two cases:

for a massless neutrino $|v_L\rangle = |v, h= -1\rangle$ and $|v_R\rangle = |v, h= +1\rangle$ can be enough because massless particles go at the speed of light (no boost can flip h)

But now we know that (at least two) neutrinos have non vanishing masses, although very small

for a completely neutral neutrino there is the possibility that neutrino and antineutrino coincide (Majorana neutrino)

Each neutrino mass eigenstate of definite helicity coincides with its own antiparticle



For a massive Majorana neutrino only two states are enough

A Majorana neutrino is identical with its charge conjugated

$$C | V > = | \overline{V} > = | V >$$

Each neutrino mass eigenstate of definite helicity coincides with its own antiparticle



Weak isospin I

$$v_{L} \Rightarrow I = 1/2, I_{3} = 1/2$$

$$v_{R} \Rightarrow I = 0, I_{3} = 0$$
Dirac Mass:
$$\nabla_{L}v_{R} + \nabla_{R}v_{L} \quad |\Delta I| = 1/2$$
Can be obtained from Higgs doublets: $v_{L}v_{R}H$
Majorana Mass:
$$v_{L}^{T}v_{L} \qquad |\Delta I| = 1$$
Non ren., dim. 5 operator: $v_{L}^{T}v_{L}HH$

$$v_{R}^{T}v_{R} \quad |\Delta I| = 0 \qquad Directly compatible with SU(2)xU(1)!$$



A very natural and appealing explanation:

v's are nearly massless because they are Majorana particles and get masses through L non conserving interactions suppressed by a large scale M ~ M_{GUT}

 $m_v \sim \frac{m^2}{M}$ m: ≤ $m_t \sim v \sim 200$ GeV M: scale of L non cons.

Note:

$$m_v \sim (\Delta m_{atm}^2)^{1/2} \sim 0.05 \text{ eV}$$

m ~ v ~ 200 GeV



M ~ 10¹⁴ - 10¹⁵ GeV

Neutrino masses are a probe of physics at M_{GUT} !





How to prove that v's are Majorana fermions? All we know from experiment on v masses strongly indicates that v's are Majorana particles and that L is not conserved (but a direct proof still does not exist).



Detection of $0\nu\beta\beta$ (neutrinoless double beta decay) would be a proof of L non conservation ($\Delta L=2$). Thus a big effort is devoted to improving present limits and possibly to find a signal.

Heidelberg-Moscow, Cuoricino-Cuore, GERDA,



$0\nu\beta\beta$ would prove that L is not conserved and v's are Majorana Also can tell degenerate, inverted or normal hierarchy

$$|m_{ee}| = c_{13}^2 [m_1 c_{12}^2 + e^{i\alpha} m_2 s_{12}^2] + m_3 e^{i\beta} s_{13}^2$$



 \oplus

Baryogenesis

$$n_{\rm B}/n_{\gamma} \sim 10^{-10}$$
, $n_{\rm B} >> n_{\rm Bbar}$

Conditions for baryogenesis: (Sacharov '67)

- B (and L) non conservation (obvious)
- C, CP non conserv'n (B-B^{bar} odd under C, CP)
- No thermal equilib'm (n=exp[μ-E/kT]; μ_B=μ_{Bbar}, m_B=m_{Bbar} by CPT

If several phases of BG exist at different scales the asymm. created by one out-of-equilib'm phase could be erased in later equilib'm phases: BG at lowest scale best

Possible epochs and mechanisms for BG:

- At the weak scale in the SM Excluded
- At the weak scale in the MSSM Disfavoured
- Near the GUT scale via Leptogenesis Very attractive



Baryogenesis by decay of heavy Majorana v's BG via Leptogenesis near the GUT scale $T \sim 10^{12\pm3}$ GeV (after inflation) Buchmuller, Yanagida, Plumacher, Ellis, Lola, Only survives if Δ (B-L) is not zero Giudice et al, Fujii et al (otherwise is washed out at T_{ew} by instantons) Main candidate: decay of lightest v_R (M~10¹² GeV) L non conserv. in v_{R} out-of-equilibrium decay: B-L excess survives at T_{ew} and gives the obs. B asymmetry. Quantitative studies confirm that the range of m_i from v oscill's is compatible with BG via (thermal) LG In particular the bound $m_i < 10^{-1} eV$ was derived for hierarchy Buchmuller, Di Bari, Plumacher; Can be relaxed for degenerate neutrinos Giudice et al; Pilaftsis et al; So fully compatible with oscill'n data!! Hambye et al
The current experimental situation on ν masses and mixings has much improved but is still incomplete

- what is the absolute scale of v masses?
- precise value of θ_{13} , shift of θ_{23} from maximal, CP viol. phase....
- pattern of spectrum (sign of Δm_{atm}^2)



- no detection of 0vββ (i.e. no proof that v's are Majorana) see-saw?
- are 3 light v's OK? (are there sterile neutrinos?)
- Different classes of models are still possible





Neutrino Masses & Mixings 2012

Guido Altarelli Universita' di Roma Tre CERN Models of v masses and mixings

An interplay of different matrices:

charged lepton diagonalisat'n

interplay of different matrices:
$$m_{\ell}' = V_{\ell}^{\dagger} m_{\ell} U_{\ell}$$

 $U_{PMNS} = U_{\ell}^{\dagger} U_{\nu}^{\dagger} m_{\ell}' = U_{\ell}^{\dagger} m_{\ell}^{\dagger} m_{\ell} U_{\ell}$
neutrino diagonalisat'n

 $m \rightarrow Rm I$

The large v mixing $O_5 = \ell^T \frac{\lambda^2}{M} \ell H H \rightarrow V_L^T m_v V_L$ versus the small q mixing can be due to the Majorana nature of v's

$$m_v' = U_v^T m_v U_v$$

See-saw $m_v = m_D^T M^{-1} m_D$ neutrino Dirac mass

neutrino Majorana mass

General remarks

• Finally not too much hierarchy is found in v masses:

 $r \sim \Delta m_{sol}^2 / \Delta m_{atm}^2 \sim 1/30$

Only a few years ago could be as small as 10⁻⁸! Precisely at 3σ : 0.025 < r < 0.039 Schwetz et al '10 or $m_{heaviest} < 0.2 - 0.7 \text{ eV}$ $m_{next} > ~8 ~10^{-3} eV$ For a hierarchical spectrum: $\frac{m_2}{m_3} \approx \sqrt{r} \approx 0.2$ Comparable to $\lambda_{\rm C} = \sin \theta_{\rm C}$: $\lambda_{\rm C} \approx 0.22 \text{ or } \sqrt{\frac{m_{\mu}}{m_{\tau}}} \approx 0.24$ Suggests the same "hierarchy" parameters for q, l, v (small powers of λ_{c}) \longrightarrow e.g. θ_{13} not too small! I now discuss some current ideas on model building We go from less to more structure

Models with little symmetry are more qualitative. Some examples:

```
Anarchy
Semianarchy
Lopsided models
U(1)<sub>FN</sub>
```

With better data the range for each mixing angle has narrowed and precise special patterns are suggested that can be reproduced by specified symmetries :

> TriBimaximal (TB), BiMaximal (BM),..... Discrete non abelian flavour groups A4, S4,....



An extreme point of view

No order for leptons -> Anarchy

In the lepton sector no symmetry, no dynamics is assumed; only chance

Hall, Murayama, Weiner'00

Boosted recently by θ_{13} near the previous bound





Anarchy and its variants can be embedded in a simple GUT context based on



Offers a simple description of hierarchies for quarks and leptons, but only orders of magnitude are predicted (large number of undetermined o(1) parameters)



Hierarchy for masses and mixings via horizontal U(1)_{FN} charges.Froggatt, Nielsen '79The simplest flavour symmetry

Principle:A generic mass term
$$\overline{R}_1 m_{12} L_2 H$$
 q_1, q_2, q_H :is forbidden by U(1)U(1) charges ofif $q_1 + q_2 + q_H$ not 0 \overline{R}_1, L_2, H

U(1) broken by vev of "flavon" field θ with U(1) charge q_{θ} = -1. If vev θ = w, and w/M= λ we get for a generic interaction: $\overline{R}_{1}m_{12}L_{2}H(\theta/M)q^{1+q^{2+qH}}$ $m_{12} -> m_{12}\epsilon^{q^{1+q^{2+qH}}}$

Hierarchy: More Δ_{charge} -> more suppression ($\epsilon = \theta/M$ small) One can have more flavons ($\epsilon, \epsilon', ...$) with different charges (>0 or <0) etc -> many versions Anarchy can be realised in SU(5) by putting all the flavour structure in T ~ 10 and not in $F^{bar} \sim 5^{bar}$

 $\begin{array}{ll} m_u \sim 10.10 & strong hierarchy \quad m_u : m_c : m_t \\ m_d \sim 5^{bar} .10 \quad \sim m_e^T & milder hierarchy \quad m_d : m_s : m_b \\ & or \quad m_e : m_\mu : m_\tau \end{array}$

Experiment supports that d, e hierarchy is roughly the square root of u hierarchy

 $m_v \sim v_L^T m_v v_L \sim 5^T .5$ or for see saw (5.1)^T (1.1) (1.5)

For example, for the simplest flavour group, $U(1)_F$

anarchy 1st fam. 2nd 3rd

$$\begin{cases}
T : (3, 2, 0) \\
F^{bar}: (0, 0, 0) \\
1 : (0, 0, 0)
\end{cases}$$



A milder ansatz - Semianarchy: no structure only in 23 Consider a matrix like $m_v \sim L^T L \sim \begin{bmatrix} \epsilon^4 & \epsilon^2 & \epsilon^2 \\ \epsilon^2 & 1 & 1 \\ \epsilon^2 & 1 & 1 \end{bmatrix}$ Note: $\theta_{13} \sim \epsilon^2 = \theta_{23} \sim 1$ with coeff.s of o(1) and det23~o(1)

["semianarchy", while $\varepsilon \sim 1$ corresponds to anarchy] After 23 and 13 rotations $m_{\nu} \sim \begin{bmatrix} \varepsilon^4 & \varepsilon^2 & 0 \\ \varepsilon^2 & \eta & 0 \\ 0 & 0 & 1 \end{bmatrix}$

Normally two masses are of o(1) or r ~1 and $\theta_{12} \sim \epsilon^2$ But if, accidentally, $\eta \sim \epsilon^2$, then r is small and θ_{12} is large.

The advantage over anarchy is that θ_{13} is naturally small and a single accident is needed to get both θ_{12} large and r small Ramond et al, Buchmuller et al, '11

SU(5)xU(1)

Recall: $m_u \sim 10\ 10$ $m_d = m_e^T \sim 5^{bar}\ 10$ $m_{vD} \sim 5^{bar}\ 1;\ M_{RR} \sim 1\ 1$

No structure for leptons No automatic det23 = 0 Automatic det23 = 0

With suitable charge assignments all relevant patterns can be obtained

1st fam. 2nd 3rd $\begin{cases} \Psi_{10}: (5, 3, 0) \\ \Psi_{5}: (2, 0, 0) \\ \Psi_{1}: (1, -1, 0) \end{cases}$ Equal 2,3 ch. for lopsided				
Model	Ψ_{10}	$\Psi_{ar{5}}$	Ψ_1	(H_u, H_d)
Anarchical (A)	(3,2,0)	(0,0,0)	(0,0,0)	(0,0)
Semi-Anarchical (SA)	(2,1,0) charg	(1,0,0) es pos	(2,1,0) itive	(0,0)
Hierarchical (<i>H_I</i>)	(6,4,0) all ch	(2,0,0) arges ((1,-1,0)	(0,0)
Hierarchical (H_{II})	(5,3,0)	(2,0,0)	(1,-1,0)	(0,0)
Inversely Hierarchical (IH_I)	(3,2,0)	(1,-1,-1)	(-1,+1,0)	(0,+1)
Inversely Hierarchical (IH_{II})	(6,4,0)	(1,-1,-1)	(-1,+1,0)	(0,+1)

Example: Normal Hierarchy



G.A., Feruglio, Masina'02 Note: not all charges positive --> det23 suppression $q(H) = 0, q(\overline{H}) = 0$ $q(\theta) = -1, q(\theta') = +1$

In first approx., with $\langle \Theta \rangle / M \sim \lambda \sim \lambda' \sim 0.35 \sim o(\lambda_{c})$ $10_{i}10_{j}$ $m_{u} \sim v_{u}$ $\begin{pmatrix} \lambda^{10} \ \lambda^{8} \ \lambda^{5} \\ \lambda^{8} \ \lambda^{6} \ \lambda^{3} \\ \lambda^{5} \ \lambda^{3} \ 1 \end{pmatrix}$, $m_{d} = m_{e}^{T} \sim v_{d}$ $\begin{pmatrix} \lambda^{7} \ \lambda^{5} \ \lambda^{5} \\ \lambda^{5} \ \lambda^{3} \ \lambda^{3} \\ \lambda^{2} \ 1 \ 1 \end{pmatrix}$ "lopsided" $\overline{5}_{i}1_{j}$ $\overline{5}_{i}1_{j}$ $m_{vD} \sim v_{u}$ $\begin{pmatrix} \lambda^{3} \ \lambda \ \lambda^{2} \\ \lambda \ \lambda' \ 1 \\ \lambda \ \lambda' \ 1 \end{pmatrix}$, $M_{RR} \sim M$ $\begin{pmatrix} \lambda^{2} \ 1 \ \lambda \\ 1 \ \lambda'^{2} \ \lambda' \\ \lambda \ \lambda' \ 1 \end{pmatrix}$

Note: coeffs. 0(1) omitted, only orders of magnitude predicted

with
$$\lambda \sim \lambda'$$

 $\overline{\mathbf{5}_{i}\mathbf{1}_{j}}$
 $\mathbf{m}_{vD} \sim \mathbf{v}_{u}$
 $\begin{pmatrix} \lambda^{3} \ \lambda \ \lambda^{2} \\ \lambda \ \lambda \ 1 \end{pmatrix}$,
 $\mathbf{M}_{RR} \sim \mathbf{M}$
 $\begin{pmatrix} \lambda^{2} \ 1 \ \lambda \\ 1 \ \lambda^{2} \ \lambda \\ \lambda \ \lambda \ 1 \end{pmatrix}$
see-saw
 $\mathbf{m}_{v} \sim \mathbf{m}_{vD}^{\mathsf{T}}\mathbf{M}_{RR}^{-1}\mathbf{m}_{vD}$
 $\mathbf{m}_{v} \sim \mathbf{v}_{u}^{2}/\mathbf{M}$
 $\begin{pmatrix} \lambda^{4} \ \lambda^{2} \ \lambda^{2} \\ \lambda^{2} \ 1 \ 1 \\ \lambda^{2} \ 1 \ 1 \end{pmatrix}$,
 $\frac{\det_{23} \sim \lambda^{2}}{2}$

The 23 subdeterminant is automatically suppressed, $\theta_{13} \sim \lambda^2$, θ_{12} , $\theta_{23} \sim 1$

This model works, in the sense that all small parameters are naturally due to various degrees of suppression. But too many free parameters!!

Examples of mechanisms for Det[23]~0 based on see-saw: $m_v \sim m_D^T M^{-1} m_D$ 1) A v_{R} is lightest and coupled to μ and τ King; Allanach; Barbieri et al..... $M \sim \begin{bmatrix} \varepsilon & 0 \\ 0 & 1 \end{bmatrix} \longrightarrow M^{-1} \sim \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 1 \end{bmatrix} \approx \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 0 \end{bmatrix}$ $m_{v} \sim \begin{bmatrix} a & b \\ c & d \end{bmatrix} \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 0 \end{bmatrix} \begin{bmatrix} a & c \\ b & d \end{bmatrix} \approx \frac{1}{\varepsilon} \begin{bmatrix} a^{2} & ac \\ ac & c^{2} \end{bmatrix}$ $m_{D} \sim \begin{vmatrix} 0 & 0 \\ v & 1 \end{vmatrix}$ 2) M generic but m_D "lopsided" Albright, Barr; GA, Feruglio, $m_{v} \sim \begin{bmatrix} 0 & x \\ 0 & 1 \end{bmatrix} \begin{bmatrix} a & b \\ b & c \end{bmatrix} \begin{bmatrix} 0 & 0 \\ v & 1 \end{bmatrix} = c \begin{bmatrix} x^{2} & x \\ v & 1 \end{bmatrix}$



Anarchy: both r and θ_{13} small by accident Semianarchy: only r small by accident H2: no accidents

GA, Feruglio, Masina'02





We now consider models with a maximum of order: based on non abelian discrete flavour groups

(a review G.A., Feruglio, Rev.Mod.Phys. 82 (2010) 2701 [ArXiv:1002.0211])

A number of "coincidences" could be hints pointing to the underlying dynamics







TB mixing is close to the data: θ_{12}, θ_{23} agree within ~ 1σ

> Schwetz et al '11 At 1σ: $\sin^2\theta_{12} = 1/3 : 0.297 - 0.329$ $\sin^2\theta_{23} = 1/2 : 0.45 - 0.58$ $\sin^2\theta_{13} = 0$: 0.008 - 0.020

A coincidence or a hint?

Called: **Tri-Bimaximal mixing**

Harrison, Perkins, Scott '02

$$v_3 = \frac{1}{\sqrt{2}}(-v_{\mu} + v_{\tau})$$
$$v_2 = \frac{1}{\sqrt{3}}(v_e + v_{\mu} + v_{\tau})$$



LQC: Lepton Quark Complementarity

 $\theta_{12} + \theta_{C} = (47.0 \pm 1.2)^{\circ} \sim \pi/4$

Suggests Bimaximal mixing corrected by diagonalisation of charged leptons

> A coincidence or a hint? Cannot be all true hints, perhaps none

Golden Ratio

Feruglio, Paris'11

$$\sin^2 \theta_{12} = \frac{1}{\sqrt{5}\phi} = \frac{2}{5+\sqrt{5}} \approx 0.276$$

A coincidence or a hint?

$$U_{GR} = \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0\\ \frac{\sin \theta_{12}}{\sqrt{2}} & -\frac{\cos \theta_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}}\\ \frac{\sin \theta_{12}}{\sqrt{2}} & -\frac{\cos \theta_{12}}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix}$$

Raidal'04.....







GA, F. Feruglio, ArXiv:1002.0211 (Review of Modern Physics)

 \bigoplus

I concentrate now on TB mixing (the most studied)

TB Mixing naturally leads to discrete flavour groups

$$U = \begin{bmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ \frac{-1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{-1}{\sqrt{2}} \\ \frac{-1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{-1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{bmatrix}$$

This is a particular rotation matrix with specified fixed angles



TB mixing

Harrison, Perkins, Scott

A simple mixing matrix compatible with all present data

$$m = \begin{pmatrix} x & y & y \\ y & x+v & y-v \\ y & y-v & x+v \end{pmatrix}$$

$$\int_{V=0}^{2} \left[\sqrt{\frac{2}{3}} \frac{1}{\sqrt{3}} 0 \right]$$

$$U = \begin{bmatrix} \sqrt{\frac{2}{3}} \frac{1}{\sqrt{3}} 0 \\ \frac{-1}{\sqrt{6}} \frac{1}{\sqrt{3}} \frac{-1}{\sqrt{2}} \\ \frac{-1}{\sqrt{6}} \frac{1}{\sqrt{3}} \frac{1}{\sqrt{2}} \end{bmatrix}$$

$$m_{v} = \frac{m_{3}}{2} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 1 & -1 \\ 0 & -1 & 1 \end{bmatrix} + \frac{m_{2}}{3} \begin{bmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{bmatrix} + \frac{m_{1}}{6} \begin{bmatrix} 4 & -2 & -2 \\ -2 & 1 & 1 \\ -2 & 1 & 1 \end{bmatrix}$$
Eigenvectors: $m_{3} \rightarrow \frac{1}{\sqrt{2}} \begin{bmatrix} 0 \\ 1 \\ -1 \end{bmatrix}$

$$m_{2} \rightarrow \frac{1}{\sqrt{3}} \begin{bmatrix} 1 \\ 1 \\ 1 \end{bmatrix}$$

$$m_{1} \rightarrow \frac{1}{\sqrt{6}} \begin{bmatrix} -2 \\ 1 \\ 1 \end{bmatrix}$$

Note: mixing angles independent of mass eigenvalues Compare with quark mixings $\lambda_c \sim (m_d/m_s)^{1/2}$

Why and how discrete groups, in particular A4, work?

TB mixing corresponds to m in the basis where charged leptons are diagonal $m = \begin{pmatrix} x & y & y \\ y & x+v & y-v \\ y & y-v & x+v \end{pmatrix}$

Crucial point 1: m is the most general matrix invariant under SmS = m and $A_{23}mA_{23} = m$ with:

$$S = \frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix} \qquad A_{23} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} \qquad \begin{array}{c} 2-3 \\ \text{symmetry} \\ \text{symmetry} \end{array}$$

Crucial point 2:

Charged lepton masses: a generic diagonal matrix is defined by invariance under T (or η T with η a phase): a

$$m_l^+ m_l = T^+ m_l^+ m_l T$$

An essential observation is that

S, T and A₂₃ are all contained in S4 S⁴=T³=(ST²)²=1 define S4

Thus S4 is the reference group for TB mixing Lam

 $m_{l} = v_{T} \frac{v_{d}}{\Lambda} \begin{pmatrix} y_{e} & 0 & 0 \\ 0 & y_{\mu} & 0 \\ 0 & 0 & y_{\tau} \end{pmatrix}$ a possible T is $T = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \omega & 0 \\ 0 & 0 & \omega^2 \end{pmatrix}$ $\omega^3 = 1 - T^3 = 1$

A4: a vast literature

A4 is the discrete group of even perm's of 4 objects. (the inv. group of a tetrahedron). It has 4!/2 = 12 elements.

A4 is a subgroup of S4 $S^2=T^3=(ST)^3=1$ define A4

A4 has 4 inequivalent irreducible representations: a triplet and 3 different singlets

3, 1, 1', 1" (promising for 3 generations!)

Ch. leptons $l \sim 3$ e^c, μ^{c} , $\tau^{c} \sim 1$, 1", 1'

Invariance under S and T is automatic in A4 while A₂₃ is not contained in A4 (2<->3 exchange is an odd perm.) But 2-3 symmetry happens in A4 if 1' and 1" symm. breaking flavons are absent or have equal VEV's [2 of S4 = 1' + 1" of A4]. **Crucial point 3:** A4 must be broken: the alignment Before SSB the model is invariant under the flavour group A4 There are flavons ϕ_T , ϕ_S , ξ ... with VEV's that break A4:

 ϕ_T breaks A4 down to G_T , the subgroup generated by 1, T, T², in the charged lepton sector ϕ_S , ξ break A4 down to G_S , the subgroup generated by 1, S, in the neutrino sector

$$\begin{array}{l} \langle \varphi_T \rangle = (v_T, 0, 0) \\ \langle \varphi_S \rangle = (v_S, v_S, v_S) \\ \langle \xi \rangle = u \ , \ \langle \tilde{\xi} \rangle = 0 \end{array} \qquad \begin{array}{l} \phi_T, \phi_S \sim \mathbf{3} \\ \xi \sim \mathbf{1} \end{array}$$

The 2-3 symmetry occurs in A4 if 1' and 1" flavons are absent

This aligment along subgroups of A4 must naturally occur in a good model



At LO TB mixing is exact $r \sim \Delta m^2_{sol} / \Delta m^2_{atm}$ The only fine-tuning needed is to account for $r^{1/2} \sim 0.2$ [In most A4 models $r^{1/2} \sim 1$ would be expected as I, $v^c \sim 3$]

When NLO corrections are included from operators of higher dimension in the superpotential each mixing angle receives generically corrections of the same order $\delta \theta_{ij} \sim o(VEV/\Lambda)$ As the maximum allowed corrections to θ_{12} (and also to θ_{23}) are numerically $o(\lambda_c^2)$, we need VEV/ $\Lambda \sim o(\lambda_c^2)$ and we typically expect:

 $\theta_{13} \sim o(\lambda_c^2)$ data are somewhat undecided

Exp: $\theta_{13} \sim$ (2.2 - 3.1) θ_{C}^2 but also (0.5 - 0.7) θ_{C}

Of course the generic prediction can be altered in ad hoc versions e.g. Lin '09 has a A4 model where $\theta_{13} \sim o(\lambda_c)$ or by allowing fine tuning

Data are not really clearcut on $q_{13} \sim o(\lambda_c^2)$ or $o(\lambda_c)$





In a typical A4 model the expansion parameter must be relatively large and some fine tuning is needed



GA, Feruglio, Merlo '12

$$\sin^2 \theta_{23} = \frac{1}{2} + \mathcal{O}(\xi)$$
$$\sin^2 \theta_{12} = \frac{1}{3} + \mathcal{O}(\xi)$$
$$\theta_{13} = \mathcal{O}(\xi)$$



Bimaximal Mixing

Now particularly interesting since θ_{13} largish

Taking the "complementarity" relation seriously:

$$\theta_{12} + \theta_{C} = (47.0 \pm 1.7)^{\circ} \sim \pi/4$$
 Raidal'04

leads to consider models that give $\theta_{12} = \pi/4$ but for corrections from the diag'tion of charged leptons

$$U_{PMNS} = U_{\ell}^{\dagger} U_{\nu}$$

Recall:
$$\lambda_{C} \approx 0.22 \text{ or } \sqrt{\frac{m_{\mu}}{m_{\tau}}} \approx 0.24$$

Normally one obtains $\theta_{12} + o(\theta_C) \sim \pi/4$ "weak compl." rather than $\theta_{12} + \theta_C \sim \pi/4$

The large deviations from BM mixing could arise from charged lepton diagonalisation

For the corrections from the charged lepton sector, typically $|\sin\theta_{13}| \sim (1 - \tan^2\theta_{12})/4\cos\delta \sim 0.15$

Needs $|\sin\theta_{13}| \sim o(\lambda_c)$ as data now suggest

 $\theta_{12} + \theta_{\rm C} \sim \pi/4$

difficult to get. Rather: $\theta_{12} + o(\theta_C) \sim \pi/4$ "weak" LQC But beware of $\mu \rightarrow e\gamma$!

GA, Feruglio, Masina Frampton et al King Antusch et al.....

$$\bar{U}_{12} = -\frac{e^{-i(\alpha_1 + \alpha_2)}}{\sqrt{2}} + \frac{s_{12}^e e^{-i\alpha_2} + s_{13}^e e^{i\delta_e}}{2}$$

$$\bar{U}_{13} = \frac{s_{12}^e e^{-i\alpha_2} - s_{13}^e e^{i\delta_e}}{\sqrt{2}}$$

$$\bar{U}_{23} = -e^{-i\alpha_2} \frac{1 + s_{23}^e e^{i\alpha_2}}{\sqrt{2}}$$

Corr.'s from s_{12}^e , s_{13}^e to U₁₂ and U₁₃ are of first order (2nd order to U₂₃) Here is a model based on S4, where BM mixing holds in 1st approximation and is then corrected by terms $o(\lambda_c)$ from the diagonalisation of charged leptons GA, Feruglio, Merlo '09 D. Meloni '11



MEG new limit on Br($\mu \rightarrow e \gamma$) < 2.4 10⁻¹² a serious constraint on SUSY models with non diagonal mass matrices at the GUT scale



Br($\mu \rightarrow e \gamma$) < 2.4 10⁻¹²: a serious constraint





when mixing angles are reproduced

GA, Feruglio, Merlo '12

From experiment: a good first approximation for quarks


In lepton sector TB or GR or BM mixing point to discrete flavor groups

What about quarks?

A problem for GUT models is how to reconcile the quark with the lepton mixings

quarks: small angles, strongly hierarchical masses abelian flavour symm. [e.g. U(1)_{FN}] neutrinos: large angles, perhaps TB or BM non abelian discrete symm. [e.g. A4]

Can be accomodated but quarks do not add any indication for discrete flavour groups



Summary on v mixing

- v mixing angles are large except for θ_{13} that is small but not too small, close to θ_c
- The measured values of ν mixing angles are compatible with TB or GR or BM
- If not a coincidence, this points to discrete flavour groups but, on the other extreme, anarchy for leptons is still a possibility
- In principle there is no contradiction between large v mixings and small q mixings, even in GUT's
- But quarks offer no new supporting evidence for discrete flavour groups
- Natural GUT models describing all fermion masses with TB or GR or BM mixing in the lepton sector are difficult to construct, in particular for SO(10)