

Lecture 2
Kink mass at one-loop

-1-

We will now discuss how quantum effects - inherent to all field theories - and the classical solution that we found, should be combined. We will do this by considering small fluctuations around the kink solution and quantizing them. We already did something similar at the end of the previous lecture, but we'll do it from a slightly different perspective now.

$$\text{Again, } S[\phi] = \int dx \mathcal{L}(\partial_\mu \phi, \phi)$$

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{g^2}{4} (\phi^2 - v^2)^2 = \frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) - V(\phi).$$

We assume that a static kink-like solution is found. We denote such a solution ϕ_k .

We write, for small fluctuations,

$$\phi(t, x) = \phi_k(x) + \chi(t, x) \quad \text{and} \quad \chi(t, x) = \sum_n a_n(t) \chi_n(x).$$

We will specify what the summation means shortly. For the time being, we assume

that $\{\chi_n(x)\}$ ~~is~~^{are} an orthonormal set of basis functions, so that $\int \chi_n(x) \chi_m(x) dx = \delta_{nm}$

We will construct these functions specifically later. Now:

$$S[\phi_k + \chi] = S[\phi_k] + \int dx \frac{\delta S}{\delta \phi(x)} \Big|_{\phi=\phi_k} \chi(x) + \int dx \times \left[\frac{1}{2} \partial_\mu \chi \partial^\mu \chi - \frac{\delta^2 V}{2 \delta \phi^2} \Big|_{\phi=\phi_k(x)} \chi^2 \right]$$

The linear term in χ vanishes

since ϕ_k is the solution of the equation of

motion; so that we obtain

-2-

$$S[\phi] = S[\phi_k] + \int dx \left[\frac{1}{2} \partial_\mu x \partial^\mu x - \frac{1}{2} \frac{\partial^2 V}{\partial \phi^2} \Big|_{\phi=\phi_k} x^2 \right].$$

We use the representation of V in terms of

"superpotential": $V(\phi) = \frac{1}{2} \left(\frac{dW}{d\phi} \right)^2$. Then

$$\frac{dV}{d\phi} = \frac{dW}{d\phi} \cdot \frac{d^2 W}{d\phi^2} \quad \text{and} \quad \frac{d^2 V}{d\phi^2} = \frac{dW}{d\phi} \frac{d^3 W}{d\phi^3} + \left(\frac{dW}{d\phi} \right)^2.$$

We therefore write

$$S[\phi] = S[\phi_k(x)] + \int dx \left[\frac{1}{2} \left(\frac{\partial x}{\partial t} \right)^2 - \frac{1}{2} x \hat{L}_2 x \right],$$

where

$$\boxed{\hat{L}_2 = -\frac{d^2}{dx^2} + \left(\frac{d^2 W}{d\phi^2} \right)^2 + \frac{dW}{d\phi} \frac{d^3 W}{d\phi^3}}, \quad \underline{\text{where}}$$

$$\underline{\phi \rightarrow \phi_k(x)}.$$

Now we want to compute the operator \hat{L}_2 .

We know from the previous lecture that

$$\phi_k(x) = \frac{m}{\sqrt{2}g} \operatorname{th}\left(\frac{mx}{2}\right), \quad \text{where } m = v g \sqrt{2} \text{ is}$$

the mass of a small
excitation in a ~~broken~~ Higgs phase

$$\text{Also } W = \frac{g}{\sqrt{2}} \left(\frac{\phi^3}{3} - v^2 \phi \right), \quad \text{so that}$$

$$\frac{dW}{d\phi} = \frac{g}{\sqrt{2}} (\phi^2 - v^2) \quad \frac{d^2 W}{d\phi^2} = \frac{g}{\sqrt{2}} \cdot 2\phi \quad \frac{d^3 W}{d\phi^3} = \frac{g}{\sqrt{2}} \cdot 2$$

As the result, substituting $\phi_k = \frac{m}{\sqrt{2}g} \operatorname{th}\left(\frac{mx}{2}\right)$
we obtain

$$\begin{aligned} \frac{dW}{d\phi} \Big|_{\phi=\phi_k} &= \frac{g}{\sqrt{2}} \left(\frac{m^2}{2g^2} \operatorname{th}^2\left(\frac{mx}{2}\right) - v^2 \right) = \frac{gv^2}{\sqrt{2}} \left(\frac{\operatorname{sh}^2\left(\frac{mx}{2}\right)}{\operatorname{ch}^2\left(\frac{mx}{2}\right)} - 1 \right) \\ &= -\frac{gv^2}{\sqrt{2} \operatorname{ch}^2\left(\frac{mx}{2}\right)} = -\frac{m^2}{2g^2 \operatorname{ch}^2\left(\frac{mx}{2}\right)} = \frac{-m^2}{2^{\frac{3}{2}} g \operatorname{ch}^2\left(\frac{mx}{2}\right)} \end{aligned}$$

$$\frac{d^2W}{dy^2} = \frac{g}{\sqrt{2}} \frac{2m}{\sqrt{2}g} \operatorname{th}\left(\frac{mx}{2}\right) = m \operatorname{th}\left(\frac{mx}{2}\right) \Rightarrow$$

$$\begin{aligned} \left(\frac{d^2W}{dy^2}\right)^2 + \frac{dW}{dy} \frac{d^3W}{dy^3} &= \frac{g}{\sqrt{2}} \cdot 2 \frac{(-m^2)}{2^{3/2} g \operatorname{ch}^2\left(\frac{mx}{2}\right)} + m^2 \operatorname{th}^2\left(\frac{mx}{2}\right) \\ &= m^2 \operatorname{th}^2\left(\frac{mx}{2}\right) - \frac{1}{2} \frac{m^2}{\operatorname{ch}^2\left(\frac{mx}{2}\right)} = m^2 - \frac{3m^2}{2 \operatorname{ch}^2\left(\frac{mx}{2}\right)} \Rightarrow \\ \boxed{\hat{L}_2 = -\frac{d^2}{dx^2} + m^2 \left(1 - \frac{3}{2 \operatorname{ch}^2\left(\frac{mx}{2}\right)}\right)} \end{aligned}$$

We would like to diagonalize the part of the action that involves the X -fields. To do this, we can find all the eigenfunctions of \hat{L}_2 and, since this is a hermitian operator, the set of its eigenfunctions is complete and orthonormal. So, we look for

$L_2 X_n(z) = \omega_n^2 X_n(z)$. As we said, we can interpret this as an eigenvalue problem in Quantum Mechanic with the potential $V(x) = m^2 \left[1 - \frac{3}{2 \operatorname{ch}^2\left(\frac{mx}{2}\right)}\right]$.

The L_2 is hermitian, the action becomes diagonal:

$$S[X] = \int dt dx \left(\sum_{n,m} \frac{1}{2} \bar{a}_n \bar{a}_m X_n(x) X_m(x) - \frac{1}{2} \sum_{n,m} \omega_m^2 a_n a_m X_n(x) X_m(x) \right) \Rightarrow$$

$$S[x] = \sum_n \int dt \left(\frac{1}{2} \dot{a}_n^2 - \frac{1}{2} \omega_n^2 a_n^2 \right). \quad \text{The} \quad -4-$$

Hamiltonian is simply obtained by changing the sign of the last term:

$$H[x] = \sum_n \left[\frac{1}{2} \dot{a}_n^2 + \frac{\omega_n^2}{2} a_n^2 \right].$$

What we have done is the standard thing we do in QFT: we mapped the problem onto a collection of "quantum oscillators" that we can now quantize. The difference is that these "oscillators" do not describe oscillations around "empty space" (vacuum) but around a classical kink-like solution.

So, the quantization condition $[a_n, \dot{a}_m] = i\delta_{nm}$. If we now write a_n, \dot{a}_n in terms of creation and annihilation operators and express $H[x]$ through them, the ~~constant~~ energy for zero-occupation number will be

$H[x_{\text{10}}] = \sum_n \frac{\omega_n}{2}$. This zero-point energy is interpreted as a quantum correction ^{the kink} ~~to~~ ^{of the} mass,

$$\boxed{\delta M_k = H[x_{\text{10}}] = \sum_n \frac{\omega_n}{2}},$$

that we would like to calculate.

There are a few things that we will have to do in connection with all this and one such thing is ^{to understand} ~~the~~ implications of the existence of the zero-mode. The zero-mode

can be obtained from the kink solution by taking the derivative:

$$x_0(x) \sim \frac{\partial \phi_k}{\partial x} = \frac{m^2}{x^{3/2} g \operatorname{ch}^2(\frac{mx}{2})}.$$

The effect of the zero mode on the kink energy is easy to understand. A slow motion of the kink can be described by a t-dependent "center": $\phi(t, x) = \phi_k(x - x_0(t))$. The energy of this solution is

$$E = \int dx \left[\frac{1}{2} \left(\frac{\partial \phi}{\partial t} \right)^2 + \frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 + V(\phi) \right].$$

$$\frac{\partial \phi}{\partial t} = \frac{d\phi_k}{dz} (-\dot{x}_0); \quad \frac{\partial \phi}{\partial x} = \frac{d\phi_k}{dz}; \quad V(\phi_k) = V(\phi_k(z))$$

Now $dx \equiv dz$ and

$$E = \int dz \left[\frac{1}{2} \left(\frac{d\phi_k}{dz} \right)^2 \dot{x}_0^2 + \frac{1}{2} \left(\frac{d\phi_k}{dz} \right)^2 + V(\phi_k) \right]$$

$M_k = \int dz \left[\frac{1}{2} \left(\frac{d\phi_k}{dz} \right)^2 + V(\phi_k) \right]$ is the kink mass.

$$\text{Hence } E = M_k + \int dz \frac{1}{2} \left(\frac{d\phi_k}{dz} \right)^2 \dot{x}_0^2$$

One can use the explicit solution ϕ_k for the kink to simplify this further: as we

saw the kink solution satisfies

$$\boxed{\frac{d\phi_k}{dx} = \frac{dW}{d\phi}}$$

and $V(\phi_k) = \frac{1}{2} \left(\frac{dW}{d\phi_k} \right)^2$. Hence,

$$\text{Hence } \int dx \left[\frac{1}{2} \left(\frac{d\phi_k}{dx} \right)^2 + \frac{1}{2} \left(\frac{dW}{d\phi_k} \right)^2 \right] = M_k \Rightarrow$$

$$\boxed{\int dx \frac{1}{2} \left(\frac{d\phi_k}{dx} \right)^2 = \frac{M_k}{2}.}$$

Hence, we find:

$$E = M_k + \int dz \frac{1}{2} \left(\frac{d\phi_k}{dz} \right)^2 \dot{x}_0^2 = M_k + \frac{M_k \dot{x}_0^2}{2}$$

As you see, the zero mode gives the kinetic energy contribution to the kink total energy. Hence, the ~~mass~~^{mass} of the kink is

$$\boxed{M_k^{1\text{-loop}} = M_k + \sum_{n \neq 0}^{\infty} \frac{1}{2} \omega_n^2}, \text{ where } \{\omega_n\} \text{ are}$$

the eigenvalues of the operator L_2 :

$$\hat{L}_2 = -\frac{d^2}{dx^2} + m^2 \left(1 - \frac{3}{2 \operatorname{ch}^2 \left(\frac{mx}{2} \right)} \right); \quad L_2 \chi_n(x) = \omega_n^2 \chi_n(x)$$

The spectrum of this operator ~~was~~ studied in the HW1; there are two ^{discrete} eigenvalues

~~discrete~~ $\omega^2 = 0$ and $\omega^2 = \frac{3m^2}{4}$ and then continuous eigenvalues from $\omega^2 = m^2$ up to $\omega^2 = \infty$

It is easy to add include discrete ~~solutions~~ eigenvalues into the sum; we need to discuss what to do with the continuum part of the spectrum. Continuous solutions are labeled

$$\text{by an index } p = \sqrt{\omega_p^2 - m^2}; \quad 0 < p < \infty.$$

The important feature of the potential

$$m^2 \left(1 - \frac{3}{2 \operatorname{ch}^2 \left(\frac{mx}{2} \right)} \right) \text{ is that it is reflection-less}$$

(see textbooks in QM).

This means that if the continuous spectrum solution is defined through a boundary condition at $x \rightarrow \infty$ as $\chi_p^+(x) = e^{ipx}$, then this solution should be $\chi_p^+(x) = e^{ipx+i\delta p}$ at $x \rightarrow -\infty$ (i.e. no reflected wave).

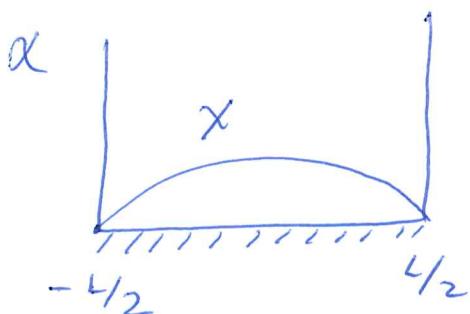
The phase δ_p is determined from the following equation

$$e^{i\delta_p} = \left(\frac{1 + ip/m}{1 - ip/m} \right) \left(\frac{1 + 2ip/m}{1 - 2ip/m} \right)$$

An independent solution describes a ~~photon~~ particle moving in the opposite direction, i.e. e^{-ipx}

It can be obtained as $\chi_p^-(x) = \chi_p^+(-x)$

Now, to have a well-defined calculation, we can put the system in a ^{large} box and require that ~~the~~ wave functions vanish at the boundaries



In general,

$$\chi(x) = A \chi_p^+(x) + B \chi_p^-(x) \Rightarrow$$

$$\chi\left(\frac{L}{2}\right) = 0 \Rightarrow A \chi_p^+\left(\frac{L}{2}\right) + B \chi_p^-\left(\frac{L}{2}\right) = 0$$

$$\chi\left(-\frac{L}{2}\right) = 0 \Rightarrow A \chi_p^+\left(-\frac{L}{2}\right) + B \chi_p^-\left(-\frac{L}{2}\right) = 0$$

$$A \chi_p^+\left(\frac{L}{2}\right) + B \chi_p^-\left(\frac{L}{2}\right) = A e^{ipL/2} + B e^{-ipL/2+i\delta_p} =$$

$$= A e^{ipL/2} + B e^{-ipL/2+i\delta_p} = \phi$$

$$0 = A \chi^+(-\frac{L}{2}) + B \chi^-_p(-\frac{L}{2}) = A e^{-ipL/2+i\delta_p} + B e^{ipL/2} \Rightarrow$$

$$\begin{pmatrix} e^{ipL/2} & e^{-ipL/2+i\delta_p} \\ e^{-ipL/2+i\delta_p} & e^{ipL/2} \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = 0 \Rightarrow$$

$$e^{ipL} - e^{-ipL+2i\delta_p} = 0 \Rightarrow \boxed{e^{ipL-i\delta_p} = \pm 1}$$

This equation of course implies that

$$pL - \delta_p = \pi n, \text{ with } n = 0, 1, \dots \text{ etc.}$$

Suppose we repeat the same calculation but with the vacuum solution; the result will be the same but for $\delta_p \equiv \phi$, so $\tilde{p}_n L = \pi n$.

The increase in the kink mass should be "renormalized", i.e. all contributions from the vacuum fluctuations $\sqrt{\text{energy density}}$ should be subtracted.

Then

$$\delta M_K = \frac{\sqrt{3}m}{4} + \sum_n \left(\frac{\omega_n}{2} - \frac{\omega_n^{\text{vac}}}{2} \right) = \frac{\sqrt{3}m}{4} + \delta \tilde{M}_K$$

$$\delta \tilde{M}_K = \sum_n \left(\frac{\omega_n}{2} - \frac{\omega_n^{\text{vac}}}{2} \right) = \frac{1}{2} \sum_n \left(\sqrt{m^2 + p_n^2} - \sqrt{m^2 + \tilde{p}_n^2} \right)$$

To calculate this sum we note that

p_n and \tilde{p}_n get close as $L \rightarrow \infty$. So we write

$$\begin{aligned} \delta \tilde{M}_K &= \frac{1}{2} \sum_n \left(\sqrt{m^2 + \tilde{p}_n^2 + (p_n^2 - \tilde{p}_n^2)} - \sqrt{m^2 + \tilde{p}_n^2} \right) \\ &= \frac{1}{2} \sum_n \left(\frac{1}{2} \frac{p_n^2 - \tilde{p}_n^2}{\sqrt{m^2 + \tilde{p}_n^2}} \right) = \frac{1}{2} \sum_n \frac{1}{2} \frac{(p_n - \tilde{p}_n)(p_n + \tilde{p}_n)}{\sqrt{m^2 + \tilde{p}_n^2}} \end{aligned}$$

$$(p_n - \tilde{p}_n)(p_n + \tilde{p}_n) \approx 2\tilde{p}_n \frac{\delta p_n}{L}. \text{ Now}$$

$$\sum_n = \frac{L}{\pi} \int dp \quad \delta \tilde{M}_K = \frac{L}{2\pi} \int_0^\infty dp \frac{1}{2} \frac{2\tilde{p}_n \delta p_n}{L \sqrt{m^2 + p^2}} \Rightarrow$$

$$\delta \tilde{M}_K = \frac{1}{2\pi} \int_0^\infty dp \delta_p \cdot \frac{\partial}{\partial p} \left[\sqrt{m^2 + p^2} \right]$$

We now integrate by parts and use the fact - that follows from the exact

expression for $e^{i\delta_p}$ on page 7 - that $\delta_p = 0$ for $p=0$ & $p=\infty$. Hence, we obtain -9-

$$\delta \tilde{M}_k = \left(-\frac{1}{2\pi}\right) \int_0^\infty dp \sqrt{m^2 + p^2} \frac{\partial \delta_p}{\partial p}.$$

We will try to pick up the logarithmic integration. To this end, calculate $\frac{d}{dp} e^{i\delta_p} = i \frac{d\delta_p}{dp} e^{i\delta_p} \Rightarrow$

$$i \frac{d\delta_p}{dp} = \frac{\frac{d}{dp} e^{i\delta_p}}{e^{i\delta_p}} = \frac{(1-ip/m)(1-2ip/m)}{(1+ip/m)(1+\frac{2ip}{m})} \times \frac{d}{dp} \left[\frac{1+ip/m}{1-ip/m} \right]$$

It is straight-forward to calculate the derivative, in the limit $p \gg m$. We find ($p \gg m$)

$$\frac{d\delta_p}{dp} = \frac{3m}{p^2}. \text{ Then, the shift in the mass}$$

$$\delta \tilde{M}_k = -\frac{1}{2m} \int_0^\infty dp p \frac{3m^2}{p^2} = -\frac{3m}{2\pi} \int_m^\infty \frac{dp}{p} = -\frac{3m}{2\pi} \ln\left(\frac{M_{uv}}{m}\right)$$

The one-loop shift of the mass becomes:

$$\delta M_k = \frac{\sqrt{3}m}{4} - \frac{3m}{2\pi} \ln\left(\frac{M_{uv}}{m}\right) + \begin{cases} \text{uncalculated} \\ \text{non-logarithmic} \\ \text{continuum terms} \end{cases}$$

The full mass of the kink at one-loop is

$$M_k^{\text{one-loop}} = M_k + \delta M_k = \frac{m^3}{3g^2} + \frac{\sqrt{3}m}{4} - \frac{3m}{2\pi} \ln\left(\frac{M_{uv}}{m}\right) + \dots$$

To get to the very final result, we need to account for the fact that ~~there is a~~ ^{there is a} one-loop renormalization of the mass parameter in QFT

$$\underline{Q} \Leftrightarrow m_R^2 = m^2 - \frac{3g^2}{2\pi} \ln\left(\frac{M_{uv}^2}{m^2}\right)$$

This allows us to absorb the $\ln \frac{M_K}{m}$ divergence into renormalized mass.

Thus we have $M_K^{\text{one-loop}} = \frac{m_K^3}{3g^2} + \frac{\sqrt{3} m_K}{4} + \dots$

The complete one-loop result (i.e. including non-logarithmic correction) reads

$$M_K = \frac{m_K^3}{3g^2} - m_K \left(\frac{3}{2\pi} - \frac{\sqrt{3}}{12} \right).$$

(Dasher, Hasslacher, Neveu)